

Geo/Astrophys. Fluid
Dynamics Seminar

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Neutron Stars: Astrophysical Superfluids

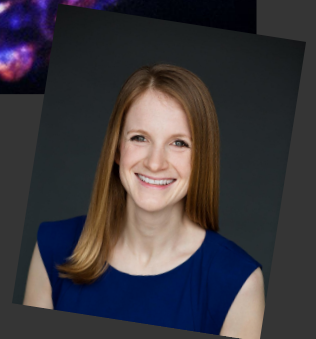
Dr. Vanessa Graber

Institute of Space Sciences
(ICE-CSIC)
Barcelona, Spain

Oct 10th, 2023



Cassiopeia A supernova remnant
(credit: NASA/CXC/SAO)



NEUTRON STARS IN A NUTSHELL

formed in
supernova
explosions

mainly
contain
neutrons

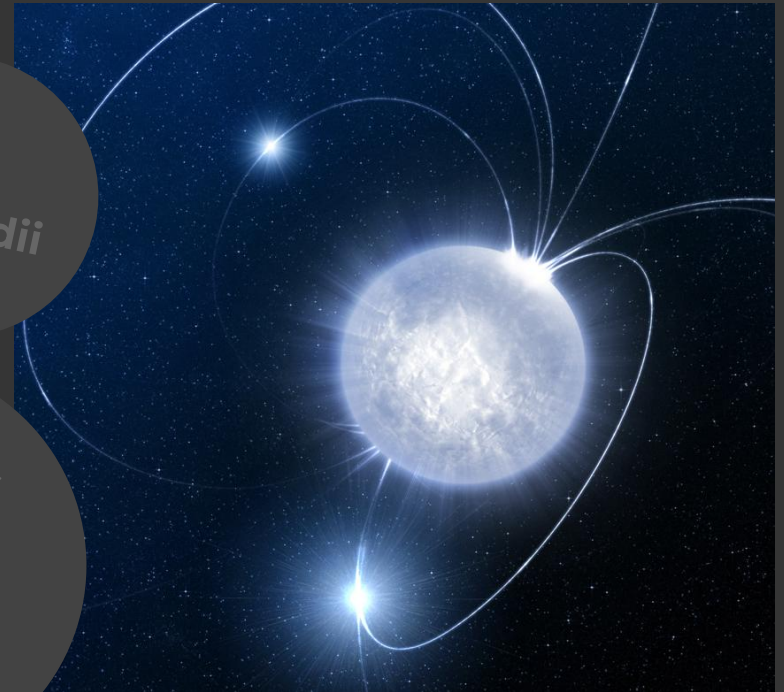
strongest
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magnets

~1-2
solar
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~10
km radii

rotate up
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densities
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credit: ESO, L. Calçada

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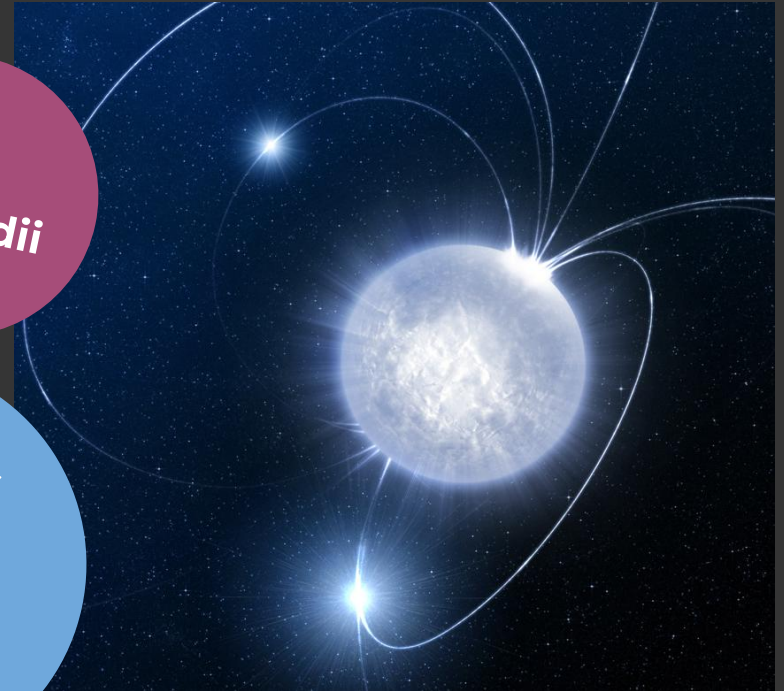
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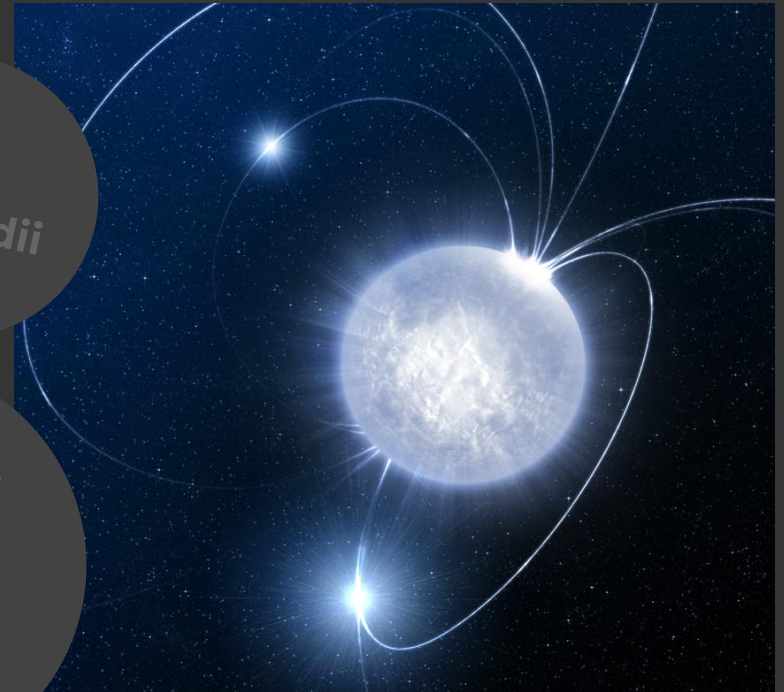
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Artist illustration of a
neutron star and its
dipolar magnetic field.



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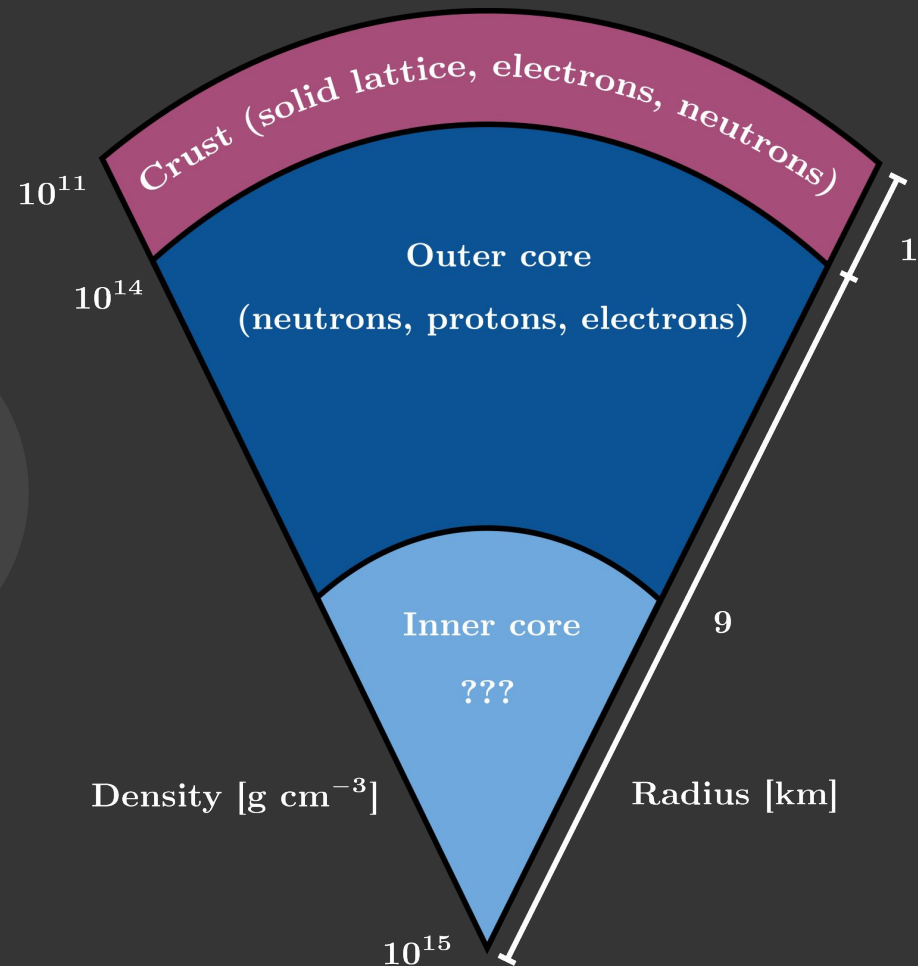
NEUTRON STAR INTERIORS

Their structure is complex and influenced by the (unknown) equation of state.

after 10^4 yrs
temperatures reach
 $\sim 10^6$ - 10^8 K

structure
is layered
much like
the Earth

decompose
into solid
crust and
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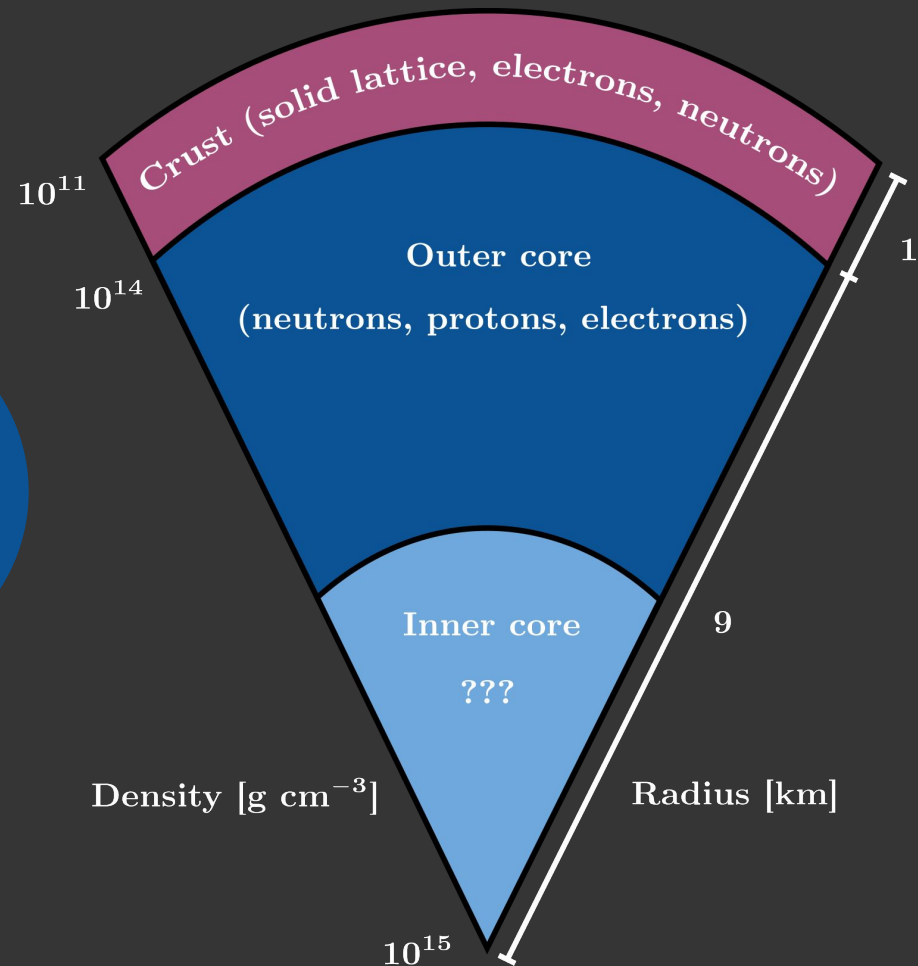
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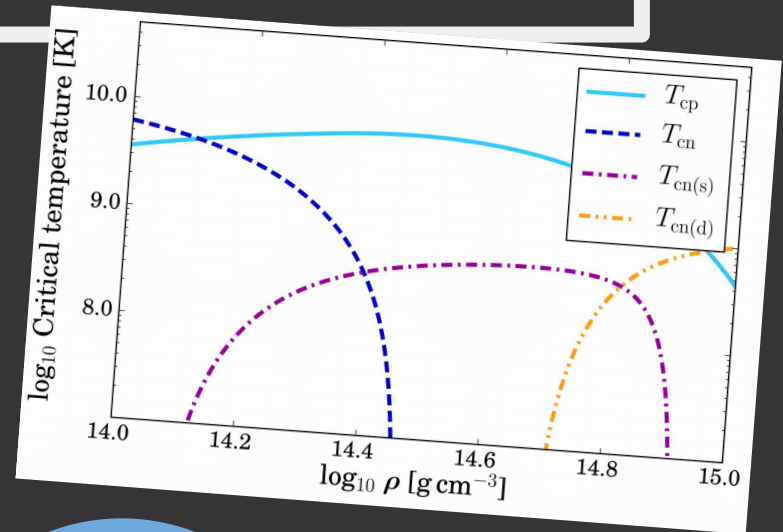
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SUPERFLUID COMPONENTS

Although neutron stars are hot compared to laboratory experiments, they are very cold in terms of their densities.



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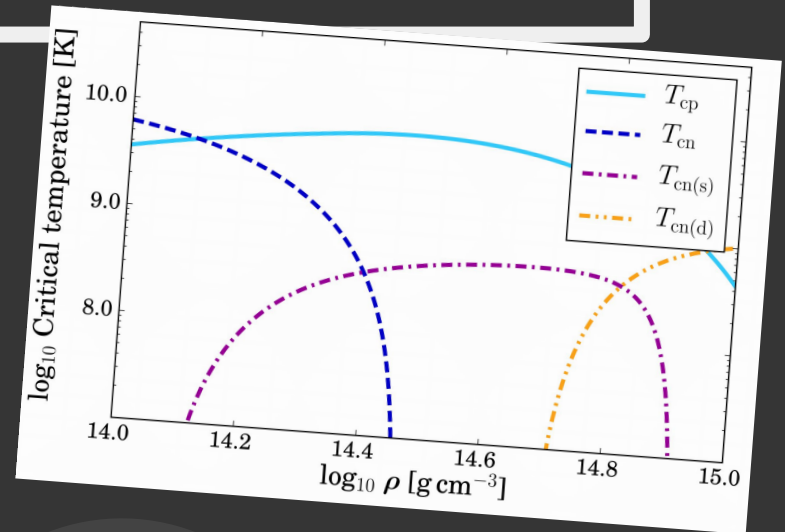
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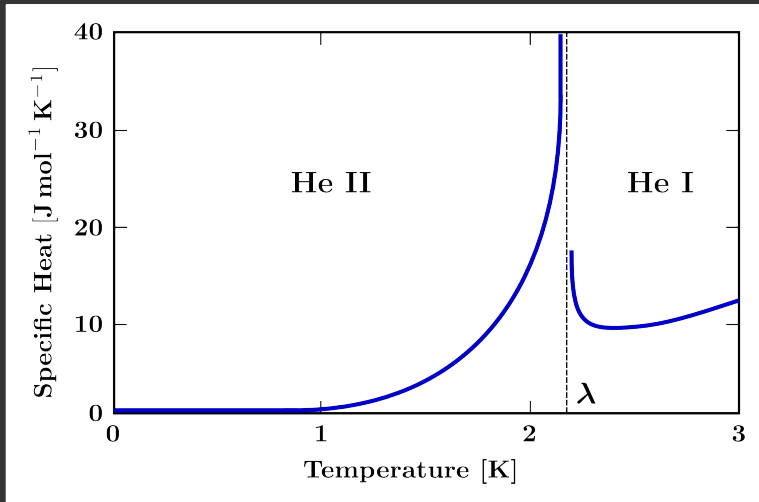
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Large numbers of particles condense into the same quantum state, which is characteristic for macroscopic quantum phenomena.

SUPERFLUID HELIUM



At low temperatures, helium-4 does not solidify but enters a new fluid phase.

heat capacity resembles the Greek letter λ

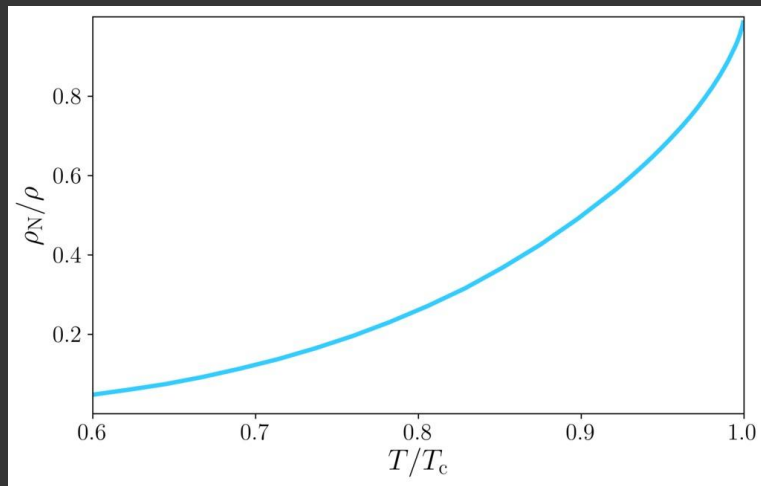
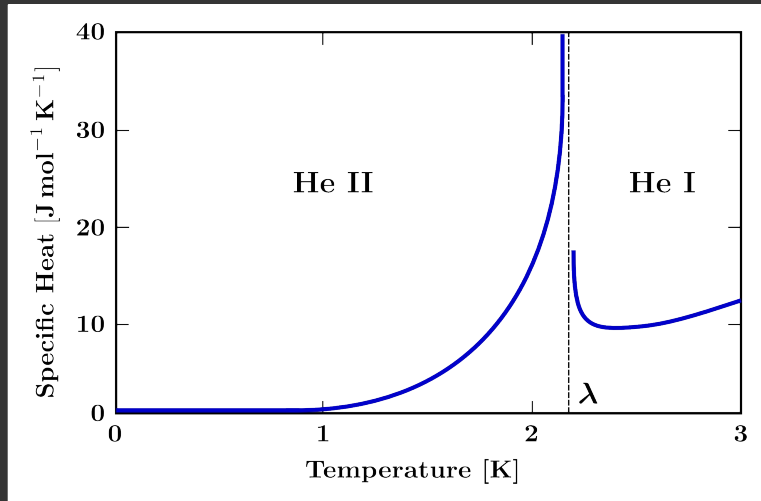
helium II behaviour explained by two-fluid model

normal component: viscous properties & heat transport

inviscid component: SF features

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QUANTUM VORTICES

Superfluids are characterised by a QM wave function $\Psi = \Psi_0 e^{i\phi}$, which satisfies the Schrödinger equation.

$$\mathbf{v}_{\text{SF}} = \hbar/m_c \nabla \phi$$

dictates

$$\boldsymbol{\omega} = \nabla \times \mathbf{v}_{\text{SF}} = 0$$

(superflow is irrotational)

each vortex carries a quantum of circulation
 $\kappa = h/m_c$

SFs rotate by forming quantised vortices

array mimics solid-body rotation:
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Credit: NOAA Photo Library

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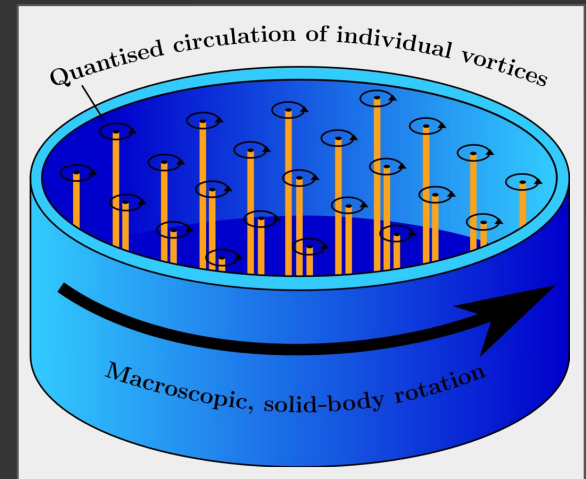
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HVBK EQUATIONS

Based on the two-fluid picture, Hall, Vinen, Bekarevich and Khalatnikov developed a vortex-averaged, hydrodynamic description of SF.

momentum / continuity equations for macroscopic velocity fields

vortices enter via tension & mutual friction

HVBK-type equations also apply to NSs: superfluid neutrons plus charged particle conglomerate

$$\rho_S \frac{D\mathbf{v}_S}{Dt} + \nabla \tilde{p}_S - \frac{\rho_S \rho_N}{2\rho} \nabla (\mathbf{v}_N - \mathbf{v}_S)^2 = \mathbf{T} + \mathbf{F}_{mf}$$

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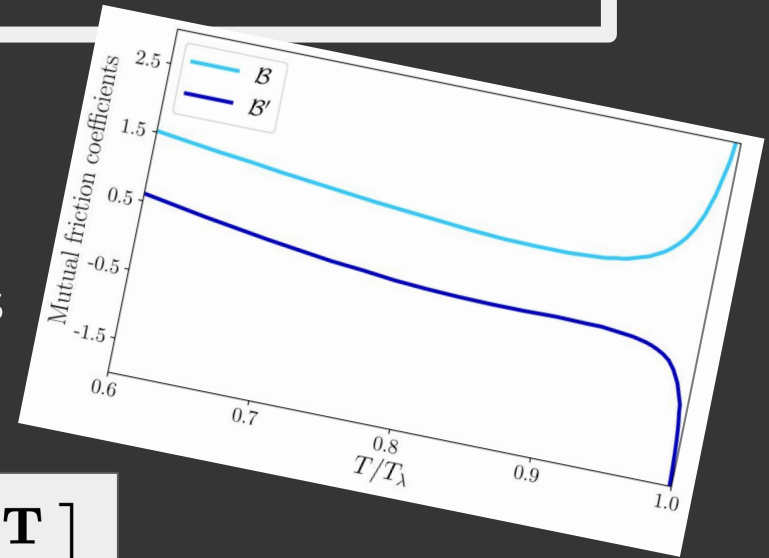
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MUTUAL FRICTION

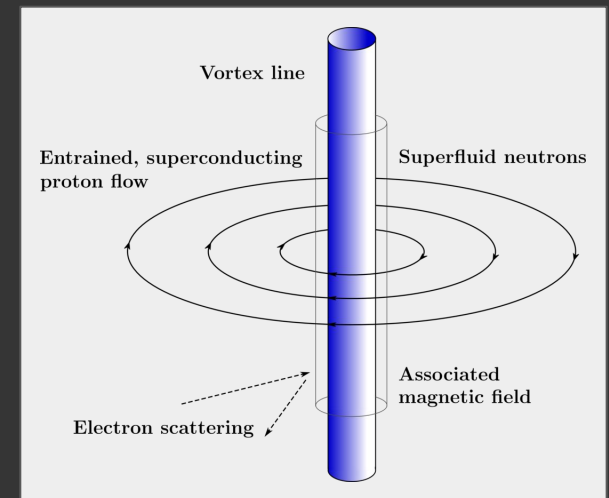
The inviscid fluid component experiences dissipation due to interactions between the vortices and the normal component.



$$\mathbf{F}_{\text{mf}} = \mathcal{B}_{\text{He}} \frac{\rho_S \rho_N}{2\rho} \hat{\boldsymbol{\omega}} \times \left[\boldsymbol{\omega} \times (\mathbf{v}_S - \mathbf{v}_N) - \frac{\mathbf{T}}{\rho_S} \right] + \mathcal{B}'_{\text{He}} \frac{\rho_S \rho_N}{2\rho} \left[\boldsymbol{\omega} \times (\mathbf{v}_S - \mathbf{v}_N) - \frac{\mathbf{T}}{\rho_S} \right]$$

two coefficients determine dissipation strength

In helium, coefficients can be directly measured. In NSs, they must be calculated.



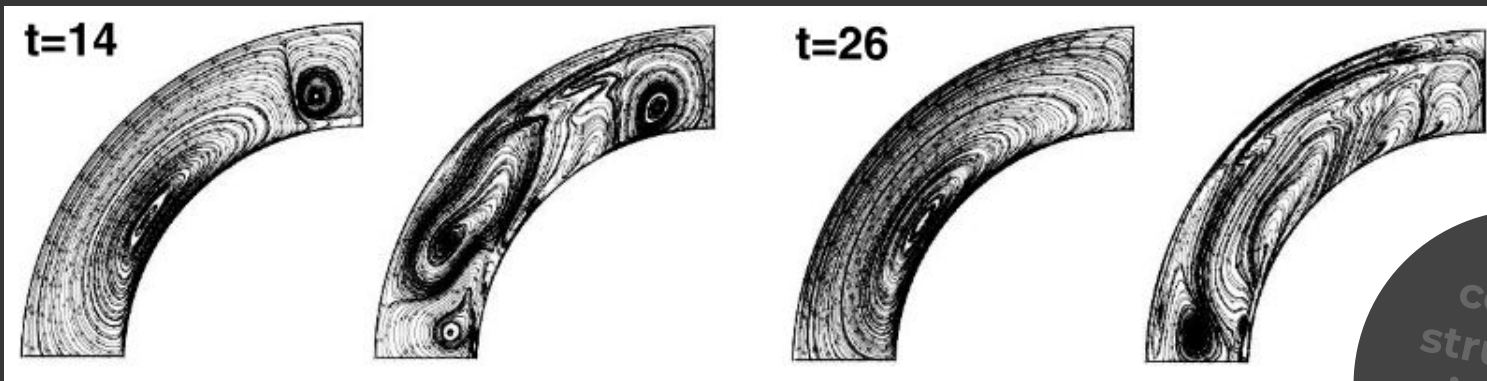
SPHERICAL NEUTRON STAR SHELL

Example: Peralta et al. (2005)
evolved the NS HVBK equations
for a rotating, spherical shell.

model
outer NS
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Streamlines for normal (left) and superfluid (right):



credit: see Fig. 1 in Peralta et al. (2005)

cell-like
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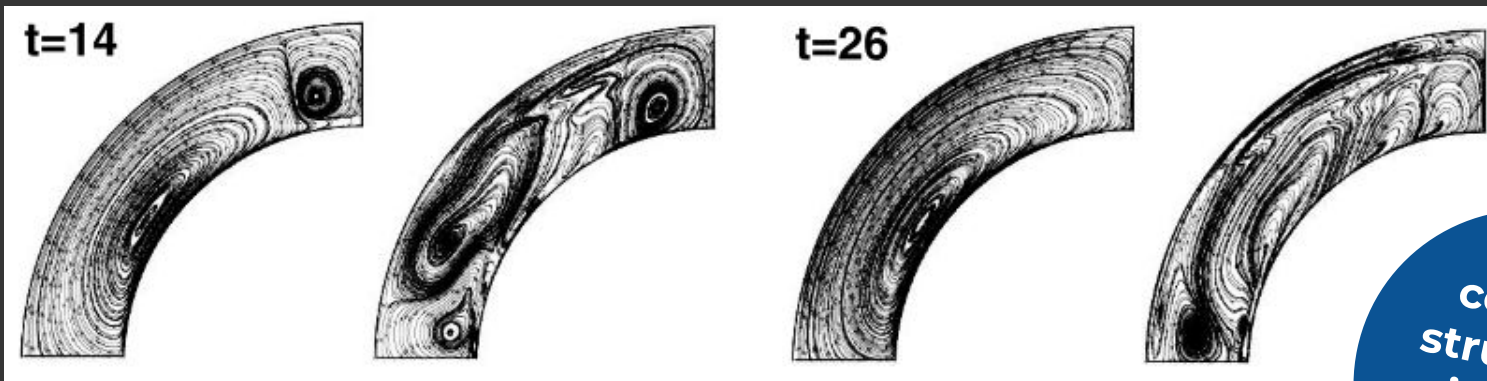
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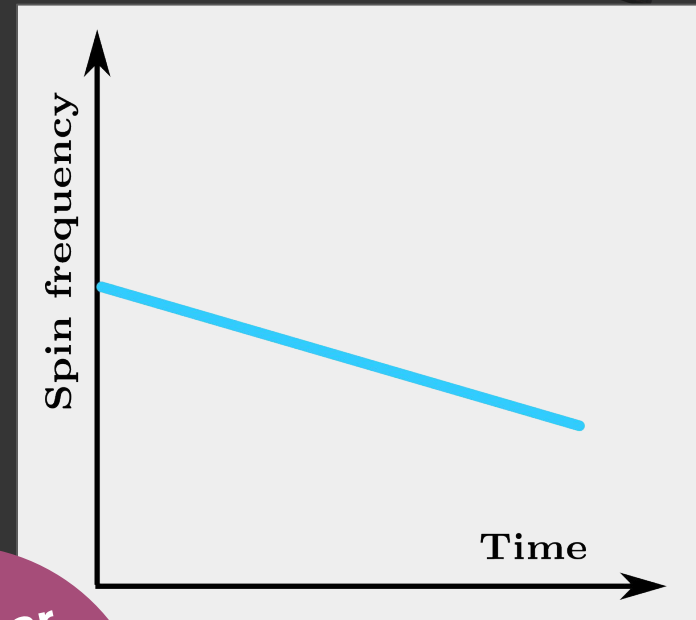


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RADIO PULSAR TIMING

Because rotation and magnetic field axes are misaligned, NSs emit radio radiation like a lighthouse.

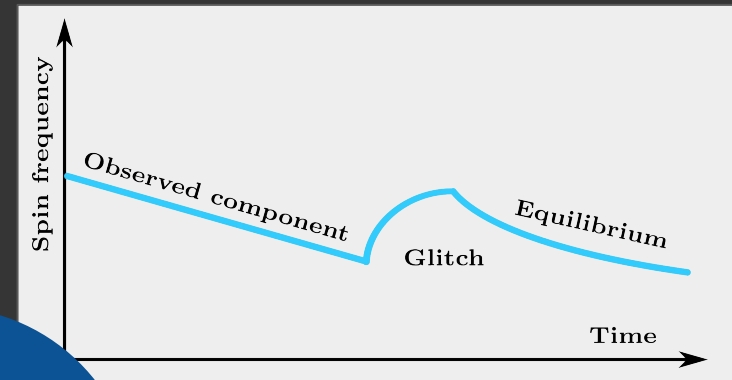


credit: J. Christiansen

time pulsar
radiation to
learn about
their
interiors

PULSAR GLITCHES

The regular spin-down of NSs can be interrupted by sudden spin-ups.

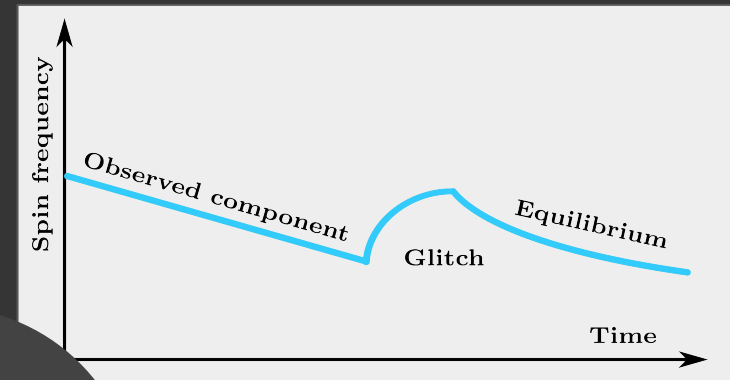


visualise
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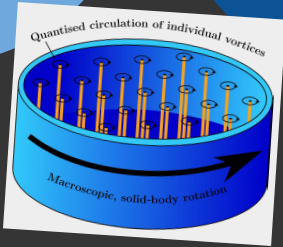
raw

MANIFESTATION OF SUPERFLUIDITY

Spin-up glitches are naturally explained in a two-component model.

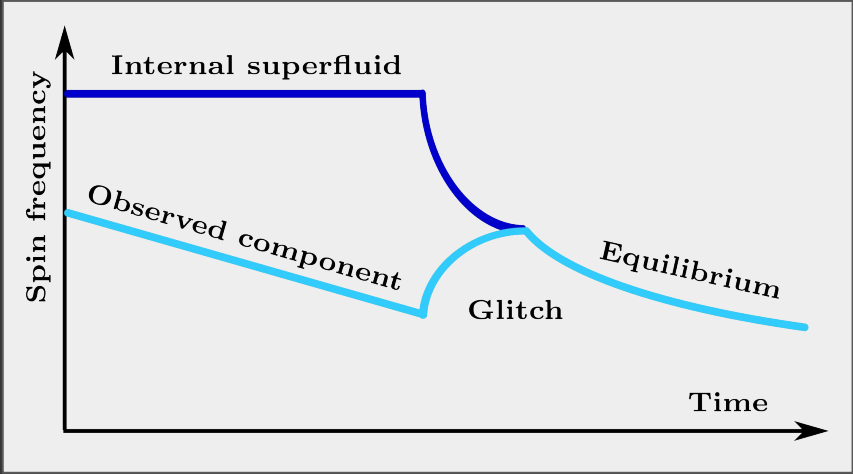
superfluid provides angular momentum reservoir

superfluid spin-down can be prevented by vortex pinning



glitches are macroscopic manifestation of quantum vortices

the shape of the glitch encodes the (hidden) internal NS physics



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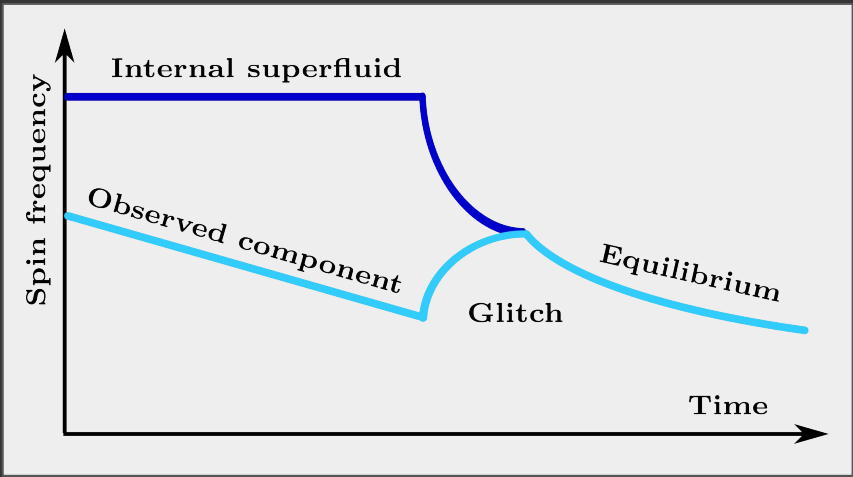
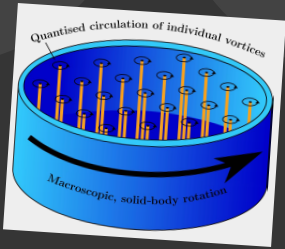
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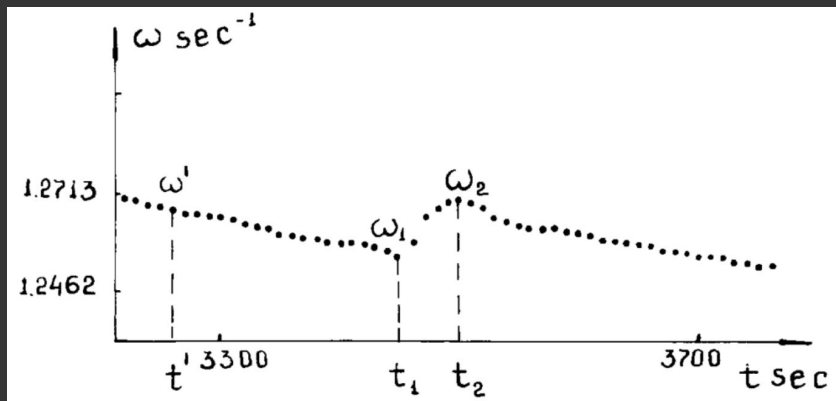
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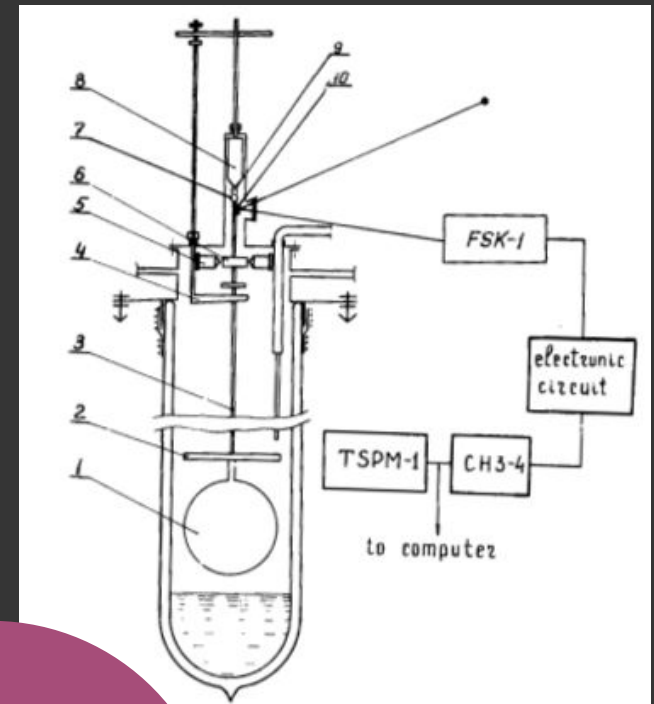


LABORATORY HELIUM GLITCHES

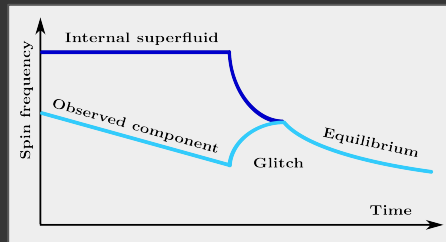
In the 1970s, Tsakadze and Tsakadze performed a systematic analysis of Helium II spin-up.



credit: Tsakadze & Tsakadze (1980)



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with their very basic set-up they might have detected one (!) glitch

NEW 3D SPIN-UP SIMULATIONS

Take a new look at numerically modelling glitch behaviour in 3D:
Solve HVBK equations with Dedalus.

$$\frac{\rho_S}{\rho} = 0.95, \quad \frac{\rho_N}{\rho} = 0.05$$

$$\mathcal{B}' = 0.90, \quad \frac{\Delta\Omega}{\Omega} = 10^{-3}$$

no
tension;
 $Re_N = 100$

start from
co-rotation
then accelerate
outer
boundary for
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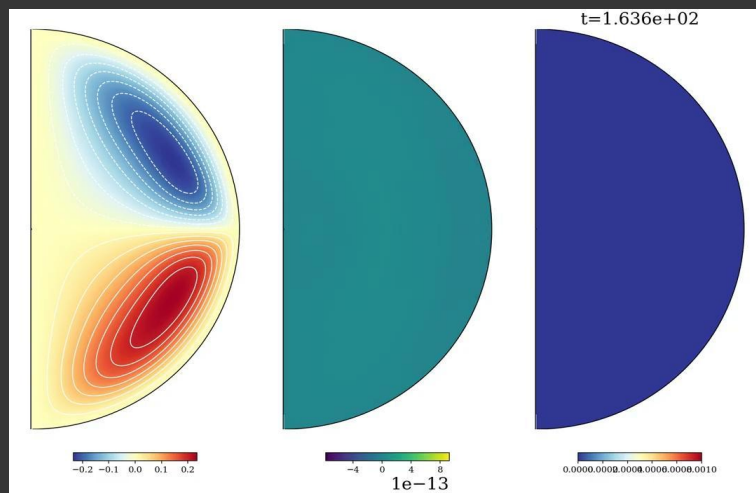
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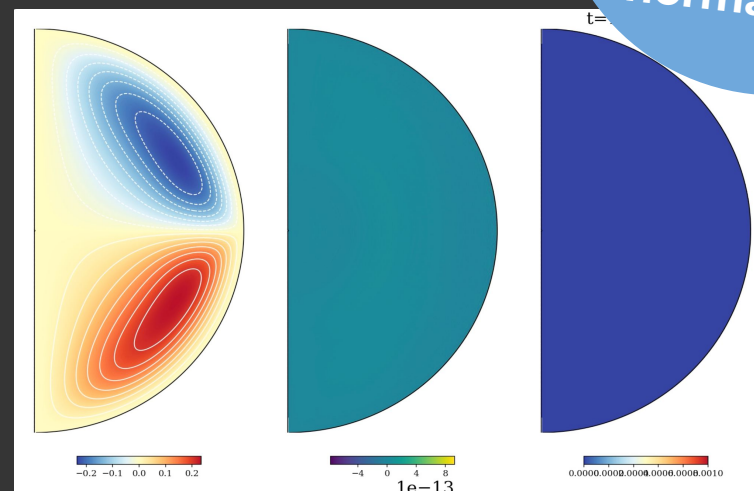
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$$\mathcal{B} = 1.5$$



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Fuentes & Graber (in prep)

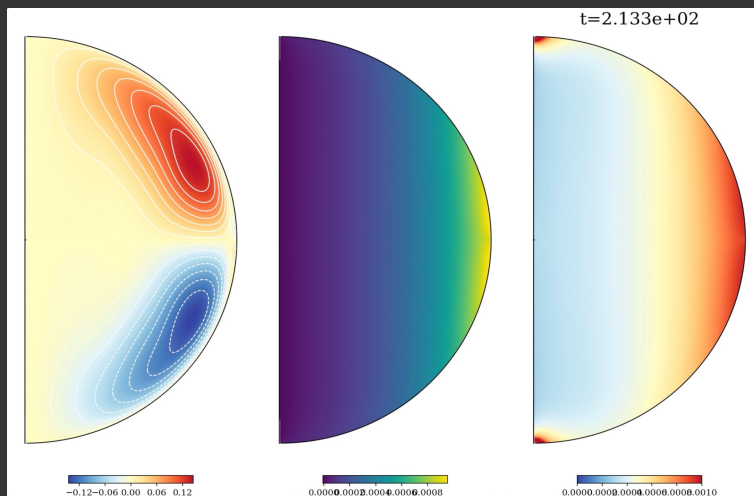
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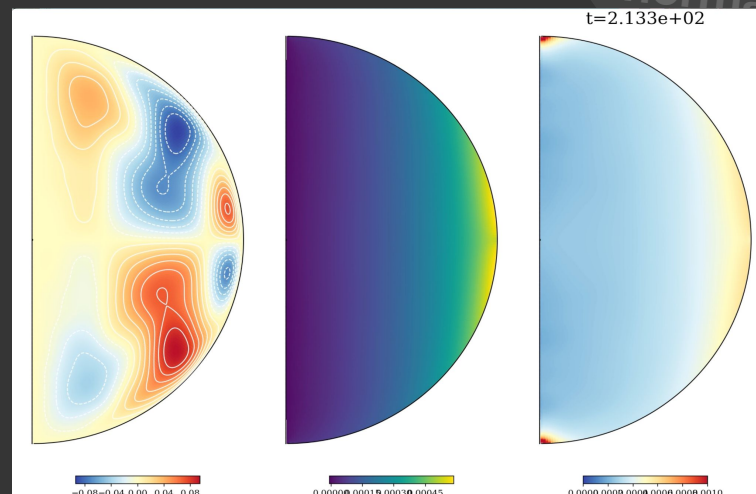
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TAKE-HOME POINTS

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